# Noisy traveling waves: effect of selection on genealogies 

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#### Abstract

For a family of models of evolving population under selection, which can be described by noisy traveling wave equations, the coalescence times along the genealogical tree scale like $\ln ^{\alpha} N$, where $N$ is the size of the population, in contrast with neutral models for which they scale like $N$. An argument relating this time scale to the diffusion constant of the noisy traveling wave leads to a prediction for $\alpha$ which agrees with our simulations. An exactly soluble case gives trees with statistics identical to those predicted for mean-field spin glasses by Parisi's theory.


Traveling wave equations such as the F-KPP equation $\partial_{t} h=\partial_{x}^{2} h+h-h^{2}[1,2,2]$ describe how a stable medium $h=1$ invades an unstable medium $h=0$. They were first introduced to study how an advantageous gene propagates through a population, $h$ being the fraction of the population with the advantageous gene. They also appear in other contexts such as disordered systems [4, 5, 6], QCD [7], reaction-diffusion [8, 9], fragmentation 10] or chemistry 11.

Traveling wave equations often represent a mean field picture where the fluctuations at the microscopic scale are ignored. The effect of these fluctuations can be represented 9, 11, 12, 13, 14] by a noise term $\left(\partial_{t} h=\right.$ $\left.\partial_{x}^{2} h+h-h^{2}+\epsilon \eta(x, t) \sqrt{h-h^{2}}\right)$. Determining quantitatively the effect of a weak noise $(\epsilon \ll 1)$ on the front position is a subject of active research. There is increasing evidence that the dynamics of the position of the front is dominated by the fluctuations near its tip 15, 16, 17 and that there is a shift in the velocity of the front 8, 18, 19, 20, 21, 22], logarithmic in the amplitude $\epsilon$ of the noise, as predicted by a simple cut-off theory [23].

In the present letter, we consider models of an evolving population under selection, which can be described by noisy traveling wave equations. Instead of focusing on the time dependence of the position of the front, we look at the problem from a different perspective: we determine how coalescence times in the genealogy depend on the size of the population. Our simulations as well as a simple argument indicate that these coalescence times scale as the inverse of the diffusion constant of the front.

We consider a population of fixed size $N$ with asexual reproduction. Each individual $i$ is characterized by a real number $x_{i}$ measuring its adequacy to the environment, and the population is fully specified by these $N$ positions $x_{i}$ 's on the adequacy axis. At each new generation, all the individuals disappear after giving birth to some offsprings. We consider the following variants:

In Model A, each individual gives birth to $k$ offsprings, and the $j$-th offspring of individual $i$ is at position $x_{i}+\epsilon_{i, j}$, where the $\epsilon_{i, j}$ are uncorrelated random numbers chosen according to some distribution $\rho(\epsilon)$. Thus, each individual inherits its parent's adequacy, and $\epsilon_{i, j}$
accounts for the effects of mutation. Then comes the selection step: out of the $k N$ new individuals, we only keep the $N$ best ones, the ones with the highest $x_{i}$ 's. Typically, we will take $k=2$ and $\rho(\epsilon)$ uniform between 0 and 1 . A similar model was proposed recently 16, 24, 25, 26] to study the evolution under competitive selection of a population of DNA molecules in vitro. The population undergoes several cycles where, in the first part of a cycle, each molecule is amplified by a fixed number $k$ with possible mutations and in the second part of the cycle selection acts by keeping only the best $1 / k$ fraction of these offsprings, defined as the molecules with the highest binding energies to a given target. In this picture, $x_{i}$ represents this binding energy.

We also investigate Model $\mathbf{A}^{\prime}$ where, instead of keeping the $N$ best individuals at each generation, we keep $N$ individuals randomly chosen among the $(3 / 2) N$ best ones. This allows us to check that our results remain unchanged under a less stringent selection.

In Model B, each individual $i$ has infinitely many offsprings, with positions distributed according to a Poisson point process of density $\psi\left(x-x_{i}\right)$ (i.e., with probability $\psi\left(x-x_{i}\right) d x$, there is an offspring of individual $i$ at position $x$ ). As in Model A, we only keep the $N$ best offsprings. Here, $\psi(\epsilon)$ is a positive function such that $\int \psi(\epsilon) d \epsilon=\infty$ (for the population not to disappear) and which decays fast enough as $\epsilon \rightarrow \infty$ to ensure that these best offsprings have finite positions. Having infinitely many offsprings at the first step is of course unrealistic, but after selection, each individual has only a finite number of offsprings. The main two reasons for considering Model B are to check the robustness of our results and to exhibit an exactly soluble case for one particular $\psi(\epsilon)$.

The genealogical tree of an evolving population can be characterized in many ways 27, 28]. Here we measure average coalescence times $\left\langle T_{p}\right\rangle$ defined as follows: $T_{p}$ is the age of the most recent common ancestor of $p$ individuals chosen at random at generation $g$, and $\left\langle T_{p}\right\rangle$ is the average of $T_{p}$ over all choices of these $p$ individuals and over all generations $g$.

In absence of selection (for example when each individual has $k$ offsprings as in model A, but with $N$ survivors


FIG. 1: Times $\left\langle T_{2}\right\rangle$ versus $N$, for different models. The scale of the $N$ axis is $\ln \ln N$. The data points are compared to several power laws of $\ln N$ shown by the straight lines.
chosen uniformly among these $k N$ offsprings), these $\left\langle T_{p}\right\rangle$ grow linearly with $N$ and their ratios take, for large $N$, the simple values (independent of $k$ ) of the Kingman coalescent [29, 30]:

$$
\begin{equation*}
\left\langle T_{p}\right\rangle \sim N, \quad \frac{\left\langle T_{3}\right\rangle}{\left\langle T_{2}\right\rangle} \rightarrow \frac{4}{3}, \quad \frac{\left\langle T_{4}\right\rangle}{\left\langle T_{2}\right\rangle} \rightarrow \frac{3}{2}, \quad \frac{\left\langle T_{p}\right\rangle}{\left\langle T_{2}\right\rangle} \rightarrow 2-\frac{2}{p} \tag{1}
\end{equation*}
$$

One goal of the present work is to show that the effect of selection changes completely Eq. (1): the time scale of these coalescence times $\left\langle T_{p}\right\rangle$ becomes

$$
\begin{equation*}
\left\langle T_{p}\right\rangle \sim[\ln N]^{\alpha} \tag{2}
\end{equation*}
$$

and the ratios are compatible with the values characterizing the Bolthausen-Sznitman coalescent 28, 31, 32]:

$$
\begin{equation*}
\frac{\left\langle T_{3}\right\rangle}{\left\langle T_{2}\right\rangle} \simeq \frac{5}{4} \quad ; \quad \frac{\left\langle T_{4}\right\rangle}{\left\langle T_{2}\right\rangle} \simeq \frac{25}{18} \tag{3}
\end{equation*}
$$

In Figs. 1 and 2 we show the results of simulations for different cases : models A and $\mathrm{A}^{\prime}$ with $k=2$ and a uniform density $\rho(\epsilon)=1$ for $0<\epsilon<1$, and model B for three different choices of the density $\psi(\epsilon): \psi_{1}(\epsilon)=\theta(-\epsilon)$, $\psi_{2}(\epsilon)=(-\epsilon)^{3} \theta(-\epsilon)$ and $\psi_{3}(\epsilon)=e^{-\epsilon}$.

Typically we simulated populations of sizes ranging from $N=10^{2}$ to $10^{5}$ over $10^{7}$ generations. We measured the times $\left\langle T_{p}\right\rangle$ by recording at each generation $g$ the age $T_{2}(i, j)$ of the most recent common ancestor of individuals $i$ and $j$. One then gets $\left\langle T_{2}\right\rangle$ by averaging $T_{2}(i, j)$ over $i, j$ and $g$. As the matrix $T_{2}(i, j)$ is ultrametric, no additional information is needed to compute the $\left\langle T_{p}\right\rangle$ : for instance, $T_{3}(i, j, k)=\max \left[T_{2}(i, j), T_{2}(i, k)\right]$. For large sizes $N$, we actually took advantage of ultrametricity by representing the matrix $T_{2}(i, j)$ as a tree: at each step, we only kept track of the current $N$ individuals and of all the most recent common ancestors of any pair of them. There are at most $N-1$ such ancestors, so both memory and execution time grow linearly with $N$, instead of $N^{2}$ if we were manipulating the full matrix.


FIG. 2: Ratios $\left\langle T_{3}\right\rangle /\left\langle T_{2}\right\rangle$ and $\left\langle T_{4}\right\rangle /\left\langle T_{2}\right\rangle$ versus $N$, for the same models (same symbols) as in Fig. $\square$ The dashed lines represent the the neutral case Eq. (11), and the dotted lines correspond to Eq. (3), i.e. model B with $\psi(\epsilon)=e^{-\epsilon}$.

In all cases except for model B with the exponential distribution, which is special as we shall see below, Fig. 1 indicates that the exponent $\alpha$ defined in Eq. (2) is in the range

$$
\begin{equation*}
2 . \leq \alpha_{\text {measured }} \leq 3 \tag{4}
\end{equation*}
$$

For model B with the exponential distribution, however, our data suggest a significantly smaller value $.75 \leq \alpha \leq 1$.

Fig. 2 shows the ratios $\left\langle T_{3}\right\rangle /\left\langle T_{2}\right\rangle$ and $\left\langle T_{4}\right\rangle /\left\langle T_{2}\right\rangle$ for the same models as in Fig. 1 In all cases, including model B with $\psi(\epsilon)=e^{-\epsilon}$, these ratios take for large $N$ values similar to Eq. (3) which differ noticeably from their values (11) in absence of selection. At present, we do not have a general theory to explain this numerical result.

Only for model B with $\psi(\epsilon)=e^{-\epsilon}$ (the special case), one can calculate these ratios exactly using a method similar to 5]. If at generation $g$, the population consists of $N$ individuals $x_{1}(g), \ldots, x_{N}(g)$, the probability of having one of their offsprings in the interval $y, y+d y$ is

$$
\sum_{i=1}^{N} \psi\left(y-x_{i}(g)\right) d y=\sum_{i=1}^{N} e^{x_{i}(g)-y} d y=e^{X_{g}-y} d y
$$

where

$$
X_{g}=\ln \left[e^{x_{1}(g)}+e^{x_{2}(g)}+\cdots+e^{x_{N}(g)}\right]
$$

So, for model B with $\psi(\epsilon)=e^{-\epsilon}$, the offsprings of the whole population are the same as those of a single effective individual at position $X_{g}$. This means that one can write $x_{i}(g+1)=X_{g}+y_{i}$, where $y_{1}, y_{2}, \ldots, y_{N}$ are the $N$ largest values of a Poisson point process on the line with exponential density. The $y_{i}$ 's are therefore distributed according to

$$
\begin{equation*}
\mathrm{P}\left(y_{N}<y_{N-1}<\cdots<y_{1}\right)=e^{-\left(y_{1}+y_{2}+\cdots+y_{N}\right)-e^{-y_{N}}} \tag{5}
\end{equation*}
$$

With this simplification, one can see that the differences $\Delta X_{g}=X_{g+1}-X_{g}$ are independent identically distributed random variables. At generation $g$, the $N$ numbers $x_{i}(g)$ form a cloud of points which does not spread in time, very much like a quantum $N$-particle bound state. This cloud has a well defined velocity $v_{N}$ and diffusion constant $D_{N}$. As the differences $\Delta X_{g}$ are independent, $v_{N}$ and $D_{N}$ are given by:

$$
v_{N}=\left\langle\Delta X_{g}\right\rangle \quad ; \quad D_{N}=\left\langle\left[\Delta X_{g}\right]^{2}\right\rangle-\left\langle\Delta X_{g}\right\rangle^{2}
$$

where the expectations are over the distribution (5) of the $y_{i}$ 's and

$$
\Delta X_{g}=X_{g+1}-X_{g}=\ln \left[e^{y_{1}}+\ldots+e^{y_{n}}\right]
$$

Calculations similar to those of [5] lead for large $N$ to

$$
\begin{equation*}
v_{N}=\ln \ln N+\frac{\ln \ln N+1}{\ln N}+\cdots ; D_{N}=\frac{\pi^{2}}{3 \ln N}+\cdots \tag{6}
\end{equation*}
$$

We now turn to computing $\left\langle T_{p}\right\rangle$ in the exponential case. If one chooses randomly $p \geq 2$ individuals at generation $g+1$, one can calculate the probability $q_{p}$ that they have the same ancestor at generation $g$

$$
\begin{equation*}
q_{p}=\left\langle\frac{\sum_{i} e^{p y_{i}}}{\left[\sum_{i} e^{y_{i}}\right]^{p}}\right\rangle \simeq \frac{1}{p-1} \frac{1}{\ln N}+\cdots \tag{7}
\end{equation*}
$$

One can also show that for large $N$ and fixed $p$, events with more than one coalescence within the $p$ individuals between two successive generations have a probability of order $(\ln N)^{-2}$ at most. Thus, for large $N$, the genealogical tree of a sample of $p$ individuals consists of single coalescence events separated by times of order $\ln N$.

From the knowledge of the $q_{p}$ 's, one can obtain for large $N$ the probability $r_{p}(k)$ that $p$ individuals at generation $g+1$ have exactly $k<p$ ancestors at generation $g$ :

$$
\begin{align*}
r_{p}(k) & =\sum_{j=0}^{k-1} \frac{(-)^{j-k+1} p!}{j!(k-1-j)!(p-k+1)!} q_{p-j}  \tag{8}\\
& =\frac{p}{(p-k)(p-k+1)} \frac{1}{\ln N}+\cdots
\end{align*}
$$

and one has

$$
\begin{equation*}
\left\langle T_{p}\right\rangle=1+\left\langle T_{p}\right\rangle+\sum_{k<p} r_{p}(k)\left[\left\langle T_{k}\right\rangle-\left\langle T_{p}\right\rangle\right] \tag{9}
\end{equation*}
$$

Using the fact that $\left\langle T_{1}\right\rangle=0$ this immediately gives

$$
\begin{equation*}
\left\langle T_{2}\right\rangle \simeq \ln N \tag{10}
\end{equation*}
$$

(in reasonable agreement with our simulations of Fig. (1) and all the ratios $\left\langle T_{n}\right\rangle /\left\langle T_{2}\right\rangle$. For $n=3$ and 4 , this gives Eq. (3) which is therefore asymptotically exact in the exponential case.

We now return to the general case. Our models are branching processes which are known to be related to
fronts of the F-KPP type 33]. Let us now see how one can associate to model B a noisy traveling wave equation (a similar calculation can be done for model A). At generation $g$ the whole population can be characterized by a function $h_{g}(x)$ which counts the fraction of individuals $i$ such that $x_{i}(g)>x$. Obviously $h_{g}(x)$ has the shape of a front $\left(h_{g}(-\infty)=1\right.$ and $\left.h_{g}(\infty)=0\right)$. From the definition of model B , one can show that $h_{g}(x)$ satisfies

$$
\begin{equation*}
h_{g+1}(x)=\min \left[1, \int h_{g}(x-\epsilon) \psi(\epsilon) d \epsilon+\frac{\eta_{g}(x)}{\sqrt{N}}\right] \tag{11}
\end{equation*}
$$

where $\eta_{g}(x)$ is some correlated noise of zero-mean with a variance equal to the integral appearing in Eq. (11). One can show as in [5] that this noise is Gaussian far from the tip of the front.

In the large $N$ limit, Eq. (11) becomes deterministic. One can look for traveling wave solutions moving at a velocity $v$. In the region where $h \ll 1$, an exponential shape $h_{g}(x) \simeq A \exp [-\gamma(x-v g)]$ is a solution if

$$
\begin{equation*}
v=v(\gamma)=\frac{1}{\gamma} \ln \left[\int e^{\gamma \epsilon} \psi(\epsilon) d \epsilon\right] \tag{12}
\end{equation*}
$$

If the front is of the F-KPP type [3], its velocity for steep enough initial conditions is the smallest one for which $\gamma$ is real. Furthermore, for finite but large $N$, one expects 15, 17, 34 a correction to this velocity of order $[\ln N]^{-2}$ and a diffusion constant scaling like $[\ln N]^{-3}$. We have checked that these predictions are compatible with our numerical simulations for all the models that we tested when $v(\gamma)$ has a minimum value (all the models of Figs. 12 except the exponential case have this property). For other choices such as $\psi(\epsilon)=\exp (-\epsilon)$, however, $v(\gamma)=\infty$ for all $\gamma$, and the front is not of the F-KPP type. It is therefore not surprising that the genealogy of the exponential case is special.

In all cases, the times $\left\langle T_{2}\right\rangle$ or $\left\langle T_{p}\right\rangle$ give the order of magnitude of the age of the most recent common ancestor of the whole population and, therefore, the time scale on which the population looses memory about its genealogy. Now, the position of this most recent common ancestor has fluctuations of order 1 , and this contributes a random shift of order 1 to the displacement of front. Therefore, the fluctuating part of the position at generation $g$ is the sum of roughly $g /\left\langle T_{2}\right\rangle$ independent random variables of order 1. Within this picture, the diffusion constant but also all the cumulants of the position would scale like

$$
\begin{equation*}
D_{N} \sim 1 /\left\langle T_{2}\right\rangle \sim 1 /\left\langle T_{p}\right\rangle \tag{13}
\end{equation*}
$$

That all cumulants have the same large $N$ dependence was indeed one of the main results of our previous work [17]. Note that Eq. (13) does hold (Eqs. (6) and (10)) in the exponential case. Assuming that it remains valid in general, we would then predict

$$
\begin{equation*}
\alpha_{\text {prediction }}=3 \tag{14}
\end{equation*}
$$

since there is now good numerical evidence 15, 34] as well as analytic arguments 17] in favor of $D_{N} \simeq[\ln N]^{-3}$ in the generic case, i.e. if $v(\gamma)$ has a finite minimum. We think that this prediction agrees with our measurements (4) up to finite size corrections: in fact, for the diffusion constant itself, it already turned out [15, 17, 34] that the large $N$ asymptotic regime was only observed for much larger systems than the ones studied here.

Beyond the fact that the time scale is logarithmic in $N$, which is not surprising for models of evolution in presence of strong selection [35, 36, 37], the times $\left\langle T_{p}\right\rangle$ are characteristic of the statistical properties of the genalogical trees of samples of a few individuals. For the exponential model, these statistics are totally specified by the fact that there is at most one single coalescence event at each generation with probability Eq. (7). Surprisingly, these coalescence probabilities (up to the factor $\ln N$ which fixes the time-scale) are the same as those which emerged from the theory of spin-glasses [28, 31, 32], so that trees, in the exponential model B here, have exactly the same statistics as the ultrametric trees of Parisi's mean-field theory of spin-glasses [38, 39]. So far we have not been able to develop a replica approach for noisy traveling waves to justify this connection. There is however some hope to do so since noisy traveling waves appear in the study of directed polymers in a random medium [5], a system for which Parisi's theory is known to be valid at the mean-field level 4].

The numerical data presented in this paper show that, for the class of models we considered, selection has drastic effects on the genealogies: the coalescence times become logarithmic in the population size (2) instead of linear and the statistics of the coalesence times are modified. The accuracy of our simulations is not sufficient to be sure that the exponent $\alpha$ and the ratios of coalescence times are universal (for all models for which $v(\gamma)$ has a finite minimum). We however gave an argument (13) which supports the conjecture (14). Of course, developing an analytical approach susceptible of proving or disproving this universality is a challenging open problem. Another open issue is whether thinking in terms of genealogies is limited to the family of selection models discussed here or could be extended to more general noisy traveling wave equations.

Lastly, it would be interesting to know what our ratios (3) would become in other models of evolution with selection such as 40, 41] and if there is a chance of estimating them from experimental data on genetic diversity.

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[1] R. A. Fisher, Annals of Eugenics 7, 355 (1937).
[2] A. Kolmogorov, I. Petrovsky, and N. Piscounov, Bull. Univ. État Moscou, A 1, 1 (1937).
[3] W. van Saarloos, Phys. Rep. 386, 29 (2003).
[4] B. Derrida and H. Spohn, J. Stat. Phys. 51, 817 (1988).
[5] É. Brunet and B. Derrida, Phys. Rev. E 70, 016106 (2004).
[6] D. Carpentier and P. Le Doussal, Nucl. Phys. B 588, 531 (2000).
[7] E. Iancu, A. H. Mueller, and S. Munier, Phys. Lett. B 606, 342 (2005).
[8] H.-P. Breuer, W. Huber, and F. Petruccione, Europhysics Letters 30, 69 (1995).
[9] C. R. Doering, C. Mueller, and P. Smereka, Physica A 325, 243 (2003).
[10] P. L. Krapivsky and S. N. Majumdar, Phys. Rev. Lett. 85, 5492 (2000).
[11] A. Lemarchand, A. Lesne, and M. Mareschal, Phys. Rev. E 51, 4457 (1995).
[12] C. Mueller and R. B. Sowers, J. Funct. Anal. 128, 439 (1995).
[13] J. Mai, I. M. Sokolov, and A. Blumen, Phys. Rev. Lett. 77, 4462 (1996).
[14] D. Panja, Phys. Rep. 393, 87 (2004).
[15] É. Brunet and B. Derrida, J. Stat. Phys. 103, 269 (2001).
[16] M. Kloster, Phys. Rev. Lett. 95, 168701 (2005).
[17] E. Brunet, B. Derrida, A. H. Mueller, and S. Munier (2006), Phys. Rev. E, in press.
[18] E. Moro, Phys. Rev. E 69, 060101(R) (2004).
[19] D. A. Kessler, Z. Ner, and L. M. Sander, Phys. Rev. E 58, 107 (1998).
[20] J. G. Colon and C. R. Doering, J. Stat. Phys. 120, 421 (2005).
[21] L. Pechenik and H. Levine, Phys. Rev. E 59, 3893 (1999).
[22] C. Escudero, Phys. Rev. E 70, 041102 (2004).
[23] É. Brunet and B. Derrida, Phys. Rev. E 56, 2597 (1997).
[24] W. Peng, U. Gerland, T. Hwa, and H. Levine, Phys. Rev. Lett. 90, 088103 (2003).
[25] R. E. Snyder, Ecol. 84, 1333 (2003).
[26] M. Kloster and C. Tang, Phys. Rev. Lett. 92, 038101 (2004).
[27] J. Schweinsberg, Elect. Journ. Prob. 5, 1 (2000).
[28] J. Pitman, Ann. Probab. 27, 1870 (1999).
[29] J. F. C. Kingman, J. Appl. Probab. 19A, 27 (1982).
[30] S. Tavaré, D. J. Balding, R. C. Griffiths, and P. Donelly, Genetics 145, 505 (1997).
[31] E. Bolthausen and A.-S. Sznitman, Com. Math. Phys. 197, 247 (1998).
[32] D. Ruelle, Com. Math. Phys 108, 225 (1987).
[33] H. P. McKean, Comm. Pure Appl. Math. 28, 323 (1975).
[34] E. Moro, Phys. Rev. E 70, 045102(R) (2004).
[35] N. L. Kaplan, R. R. Hudson, and C. H. Langley, Genetics 123, 887 (1989).
[36] R. Durrett and J. Schweinsberg, Theor. Population Biol. 66, 129 (2004).
[37] R. Durrett and J. Schweinsberg, Stoch. Proc. Appl. 115, 1628 (2005).
[38] G. Parisi, J. Phys. A 13, 1101 (1980).
[39] M. Mézard, G. Parisi, N. Sourlas, G. Toulouse, and M. A. Virasoro, Journal de Physique 45, 843 (1984).
[40] L. Tsimring, H. Levine, and D. A. Kessler, Phys. Rev. Lett. 76, 4440 (1996).
[41] D. A. Kessler, H. Levine, D. Ridgway, and L. Tsimring, J. Stat. Phys. 87, 519 (1997).

